### **OMI NO<sub>2</sub> Validation at JPL/Table Mountain Facility**



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#### **Approach**

- Direct solar absorption using the FTUVS interferometer (UV-vis-NIR) at Table Mountain Facility, north of Los Angeles, California
- Spectral resolution 0.0013 nm resolves NO<sub>2</sub> vibronic features
- Doppler-shifted spectra from east & west solar limbs are ratioed to remove Fraunhofer lines no high-sun reference spectra needed.
- Method retrieves diurnal, absolute NO<sub>2</sub> column abundance
- A "calibration" method for other techniques.

#### <u>Measurements</u>

- Data acquired twice per week, on average (FTUVS is time-shared with other Aura validation activities).
- OMI NO<sub>2</sub> column abundances are retrieved from the AVDC archive.
- Comparisons between FTUVS and OMI are sorted according to distance between Table Mtn. and the center of the closest OMI footprint.

NO<sub>2</sub> column abundances are derived from fits within microwindows that contain 10-20 NO<sub>2</sub> rotational lines.

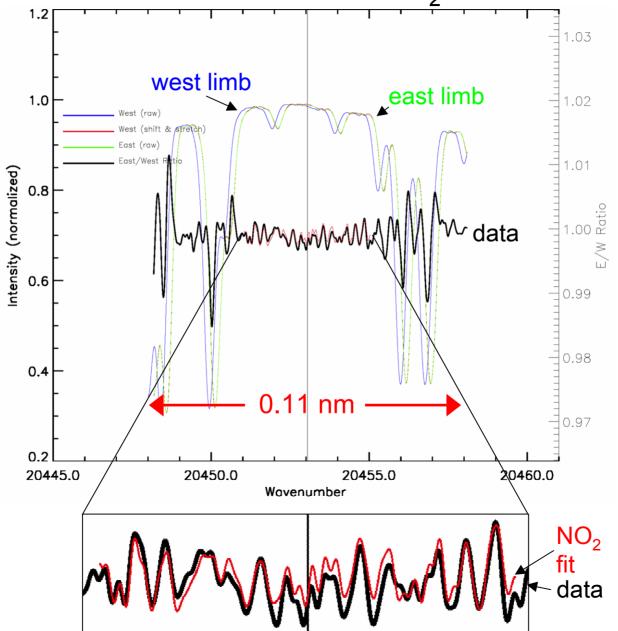
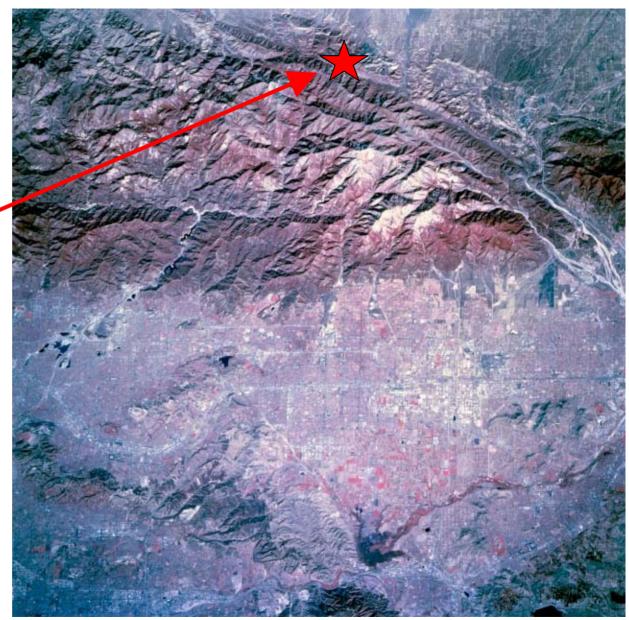
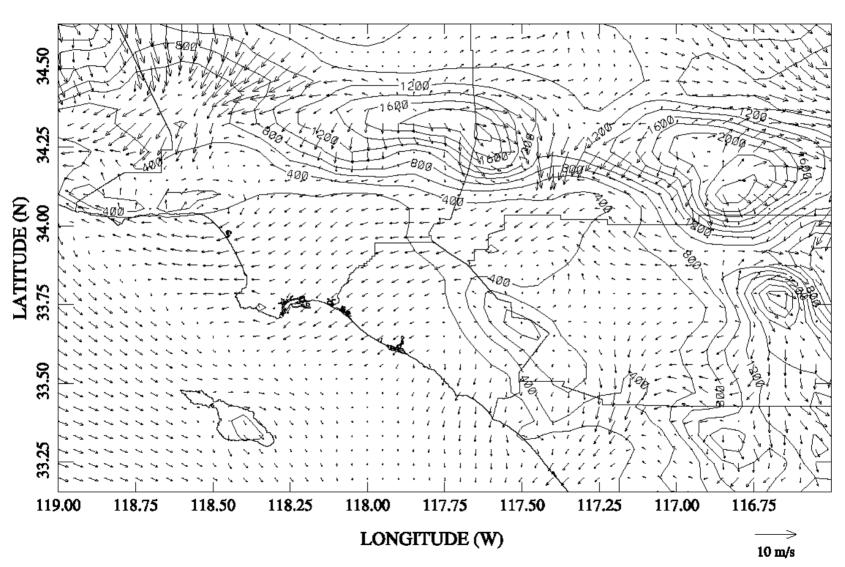


Table Mtn.
Facility is
located at
7500 ft. on the
northward
facing slopes
of the San
Gabriel Mtns.



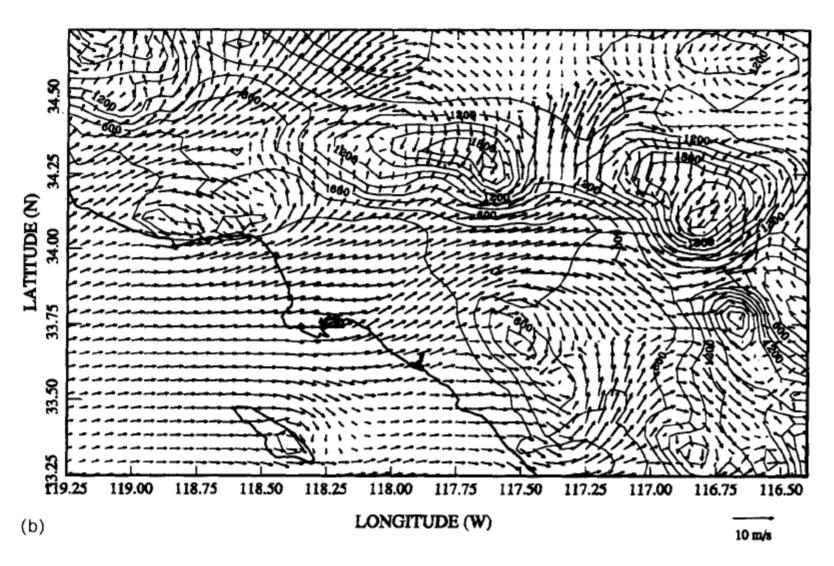


# UCLA SMOG model, 0400 PST

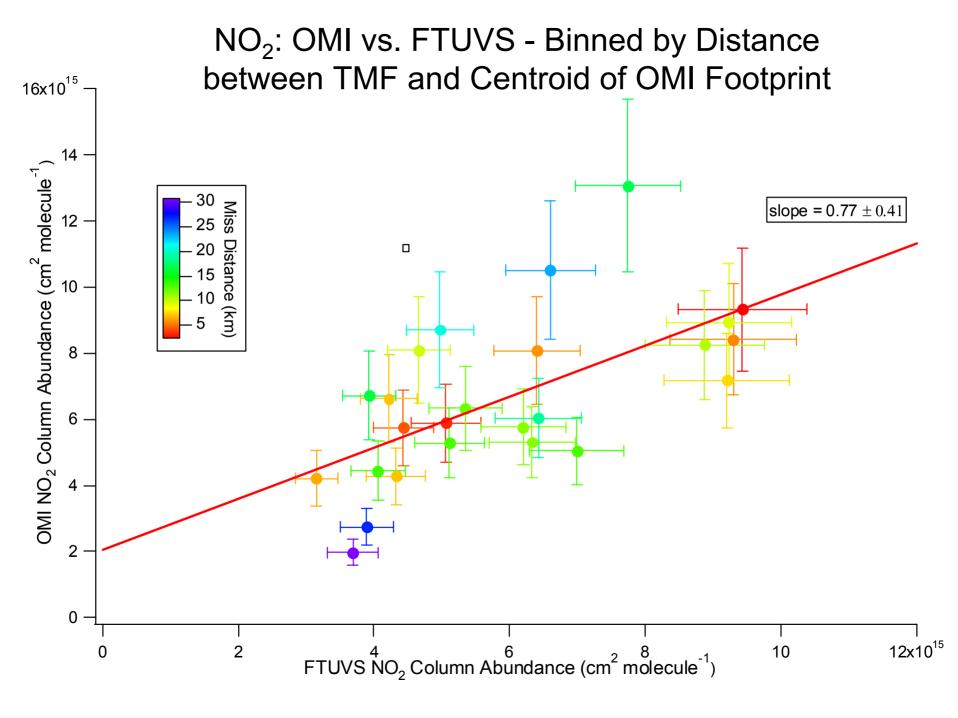


Lu et al. (2003)

# UCLA SMOG model, 1700 PST

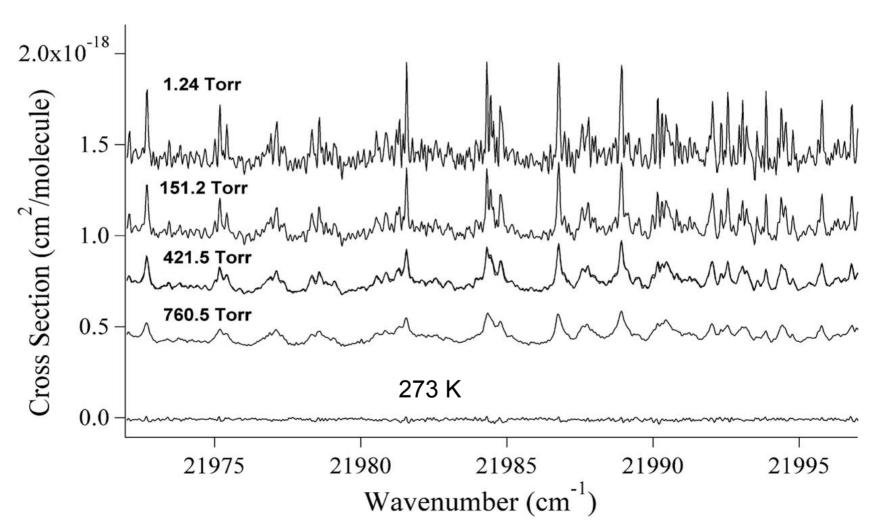


Lu and Turco (1995)

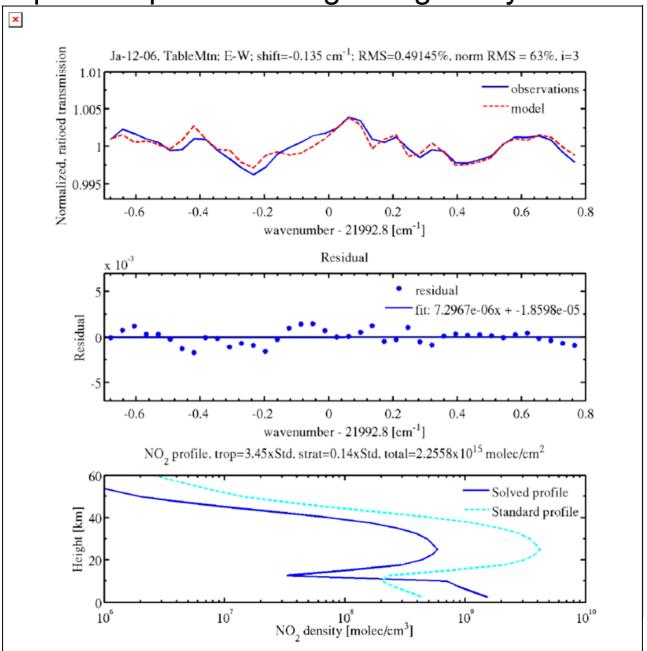


### Pressure Broadening in NO<sub>2</sub> Reference Spectra

Comparison of NO<sub>2</sub> spectral features recorded at 0.05 cm<sup>-1</sup> resolution.



Optimal Spectral Fitting Using 2-Layer Model



## Summary and Future Work

- The correlation between OMI and FTUVS NO<sub>2</sub> total column is reasonable – OMI lies 20-30% below FTUVS.
- Complex topography and circulation require that the OMI footprint centroid should be less than 10 km from TMF.
- Intercomparison campaign for NO<sub>2</sub> and O<sub>3</sub> at TMF with the WSU MF-DOAS and GSFC direct sun instruments in May '07. Other groups welcome.
- FTUVS NO<sub>2</sub> 2-level retrieval shows promise development is continuing.